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Plate Tectonics and the Archean Earth

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Abstract

If we accept that a critical condition for plate tectonics is the creation and maintenance of a global network of narrow boundaries separating multiple plates, then to argue for plate tectonics during the Archean requires more than a local record of subduction. A case is made for plate tectonics back to the early Paleoproterozoic, when a cycle of breakup and collision led to formation of the supercontinent Columbia, and bimodal metamorphism is registered globally. Before this, less preserved crust and survivorship bias become greater concerns, and the geological record may yield only a lower limit on the emergence of plate tectonics. Higher mantle temperature in the Archean precluded or limited stable subduction, requiring a transition to plate tectonics from another tectonic mode. This transition is recorded by changes in geochemical proxies and interpreted based on numerical modeling. Improved understanding of the secular evolution of temperature and water in the mantle is a key target for future research.

- Higher mantle temperature in the Archean precluded or limited stable subduction, requiring a transition to plate tectonics from another tectonic mode.

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- Plate tectonics can be demonstrated on Earth since the early Paleoproterozoic (since c. 2.2 Ga), but before the Proterozoic Earth's tectonic mode remains ambiguous.
- The Mesoarchean to early Paleoproterozoic (3.2–2.3 Ga) represents a period of transition from an early tectonic mode (stagnant or sluggish lid) to plate tectonics.
- The development of a global network of narrow boundaries separating multiple plates could have been kick-started by plume-induced subduction.

1. INTRODUCTION

After the Moon-forming impact, Earth had a core of similar size to that of the present day, a silicate mantle and crust of largely unknown compositional and mineralogical structure, and temperate surface conditions (Carlson et al. 2014). Sometime thereafter plate tectonics emerged. What tectonic mode preceded plate tectonics, how and when plate tectonics emerged, and how plate tectonics evolved are fundamental questions in geology. However, they are questions that have proven difficult to answer.

The theory of plate tectonics developed from geophysical studies of the ocean basins and was confirmed by interpretations of contemporary earthquake activity (Oreskes 2002). Plate tectonics is widely considered to have emerged on Earth sometime during the late Archean or early Proterozoic (e.g., Brown 2006, Cawood et al. 2006, Sizova et al. 2010, van Hunen & Moya 2012, Hawkesworth et al. 2016, Brown & Johnson 2018, Cawood et al. 2018, Hawkesworth & Brown 2018). However, some argue for plate tectonics as early as the Hadean (e.g., Harrison 2009), or at least for subduction at that time (Turner et al. 2014), although the style may have been different (Foley 2018). Others argue that the modern episode of plate tectonics began in the Neoproterozoic, requiring one (or more) alternative tectonic mode before this time (e.g., Stern 2018) (see the sidebar titled Major Divisions of Geologic Time).

If a critical condition for plate tectonics is the creation and maintenance of a global network of narrow plate boundaries separating multiple plates of rigid lithosphere (Bercovici et al. 2015, Mallard et al. 2016, Lenardic 2018), then regions of localized subduction that were not part of a network of plate boundaries, such as may exist on Venus (Davaille et al. 2017), are not a priori evidence for plate tectonics (Smrekar et al. 2018). At the outset, we emphasize this fundamental distinction between plate tectonics and subduction—the former, which necessarily includes subduction as a component, is of global (planetary) extent (**Figure 1**), whereas the latter may be only a regional feature on a planet without plate tectonics.

MAJOR DIVISIONS OF GEOLOGIC TIME

In chronostratigraphy, the Precambrian is the period of time from the formation of Earth around 4,567 millions of years (Ma) to the first appearance of hard-bodied fossils in the rock record about 541 Ma (the base of the Cambrian Period at the beginning of the Phanerozoic Eon). The Precambrian is subdivided into the Hadean (pre-4,000 Ma), the Archean (from 4,000 to 2,500 Ma), and the Proterozoic (from 2,500 to 541 Ma) eons. In turn, the Archean and Proterozoic eons are further subdivided into the Eoarchean (from 4,000 to 3,600 Ma), Paleoarchean (from 3,600 to 3,200 Ma), Mesoarchean (from 3,200 to 2,800 Ma), and Neoarchean (from 2,800 to 2,500 Ma) eras, and the Paleoproterozoic (from 2,500 to 1,600 Ma), Mesoproterozoic (from 1,600 to 1,000 Ma), and Neoproterozoic (from 1,000 to 541 Ma) eras, respectively.

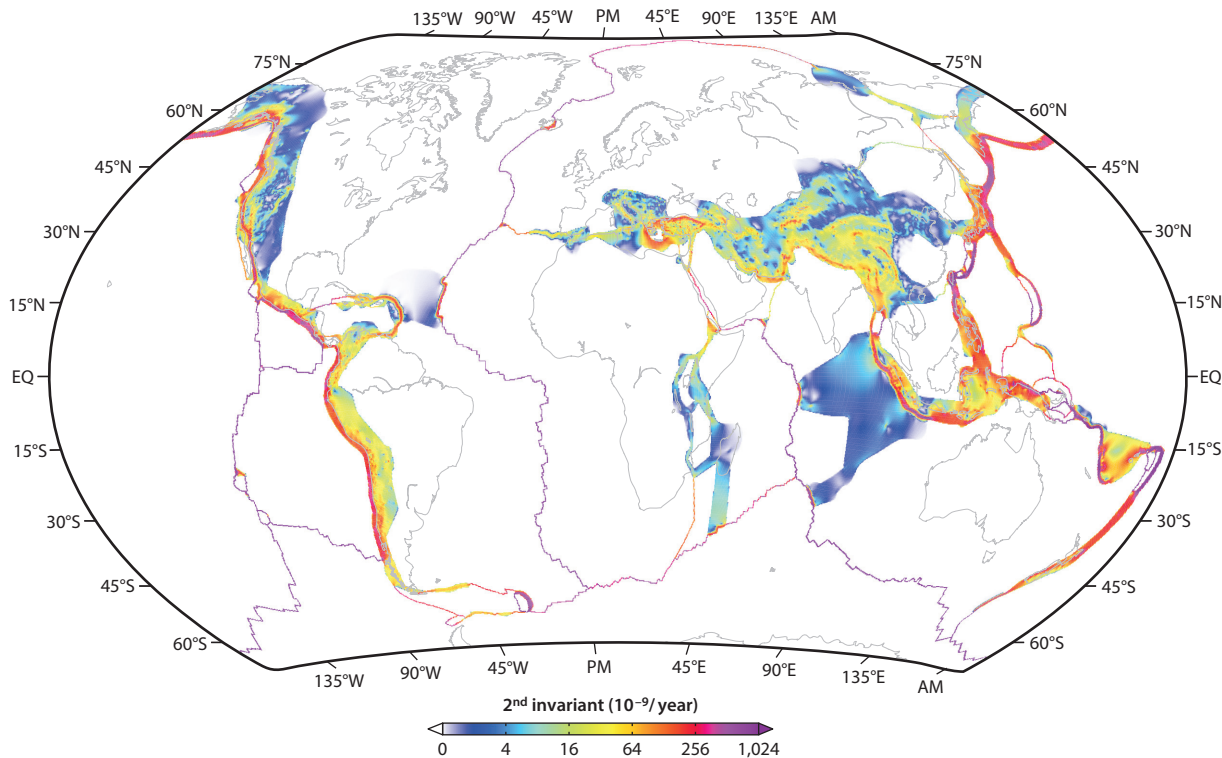


Figure 1

Map of present-day plate boundary deformation zones. White areas were assumed to be rigid plates, and the rigid body rotation of these plates was imposed as a boundary condition when solving for plate boundary strain rates from geodetic velocities. Figure adapted from Kreemer et al. (2014).

Before the Proterozoic, whether plate tectonics operated on Earth is contentious for two main reasons. First, Earth's mantle has likely been cooling since the Mesoarchean (Labrosse & Jaupart 2007, Korenaga 2013). Therefore, there may be differences in the geochemical, petrological, structural, and metamorphic features preserved in the Archean crust compared to those based on studies of contemporary convergent plate boundaries due to plate tectonics (van Hunen & Moyen 2012). Second, the amount of preserved continental lithosphere diminishes back through time and is particularly scarce before the Neoarchean. Given the likelihood of a strong survivorship bias, cratons may be atypical of the lithosphere that existed in the Neoarchean (Kamber & Tomlinson 2019). In short, the limited Archean rock record has made it difficult to demonstrate the existence of a global network of narrow plate boundaries separating multiple plates, such that it may not be possible to determine if the Archean Earth had plate tectonics (Lenardic 2018).

After we consider what potential tectonic modes could have operated on Earth and discuss the transitions between them, our strategy is to identify a lower limit to the emergence of plate tectonics on Earth from evidence preserved in the geological record. To do this we identify the two essential characteristics of plate tectonics—large horizontal displacements between continental fragments and bimodal metamorphism—and trace them back in time. These demonstrate the probable existence of a global network of plate boundaries and multiple plates since the early Paleoproterozoic, providing a minimum age for the emergence of plate tectonics. Then, we discuss the use of geochemical proxies in crustal rocks. During the Archean, some of these have been

applied to infer a local tectonic mode, such as subduction; in these cases, it may be inappropriate to infer global changes from the spatially limited data. However, for those geochemical proxies that are likely to record a global signature, this approach may provide an upper limit for the first imprint of plate tectonics in the geological record.

2. TECTONIC MODES ON PLANETS

2.1. Potential Tectonic Modes

According to Lenardic (2018), tectonic modes that may operate on terrestrial planets are (*a*) an active (mobile) lid mode, either with narrow plate boundaries or with distributed deformation; (*b*) a sluggish lid mode, either with ridges or sublithospheric convection-driven tectonics; (*c*) an episodic (transient) mode; and (*d*) a stagnant lid mode, involving heat pipe volcanism or plume tectonics. Any of the last three is a potential precursor to plate tectonics.

Plate tectonics is an active lid mode because the cold upper boundary layer of mantle convection (the lithosphere) participates in convective overturn via formation of weak zones that define narrow boundaries (Bercovici et al. 2015, Crameri & Tackley 2016), allowing it to sink (subduct) back into the interior and drive deformation (orogenesis) at convergent boundaries. If the cold upper boundary layer does not take part in convective overturn, as in a stagnant lid mode, the rate of heat loss will be slower and the style of deformation different (Korenaga 2013, Lenardic 2018). In this mode, a plutonic squishy lid is compatible with the persistence of melt-bearing lower crust for hundreds of millions of years (Taylor et al. 2020, Vanderhaeghe et al. 2019), whereas a heat pipe regime generates crust that is too cold (Sizova et al. 2015, Rozel et al. 2017). An episodic mode reflects alternating states of active and stagnant lid behavior (O'Neill et al. 2016, Lenardic 2018). In a sluggish lid mode the velocity of the lithosphere is lower than that of the convecting mantle, and mobility is driven by convective traction on the base of the lithosphere rather than by subduction (Foley 2018). Alternatively, a mixed mode, in which a poorly mobile lid domain coexists with a pseudoplate tectonics domain, could have existed between a Hadean stagnant lid and modern plate tectonics (Capitanio et al. 2019).

2.2. Mantle Temperature and Tectonic Mode

To compare mantle temperature in space and time, McKenzie & Bickle (1988) proposed the concept of mantle potential temperature (T_p)—the temperature the mantle would have at the surface if extrapolated along an adiabat without melting. According to Herzberg et al. (2007), the present-day range of T_p for mid-ocean ridge basalt (MORB) is $\sim 120^\circ\text{C}$ (1280–1400°C). However, a warmer mantle in the past, with T_p varying from place to place as it does today, may have resulted in an alternative tectonic mode during part or all of the Archean.

In principle, secular cooling of Earth may be deduced from compositional variations of mantle melts through time and from conditions at the start of mantle convection in the Hadean, when T_p may have been $\sim 200^\circ\text{C}$ warmer than at present ($\Delta T_p = +200^\circ\text{C}$; Labrosse & Jaupart 2007). Calculations suggest T_p peaked at about 3 Ga but cannot be extrapolated further back in time (Labrosse & Jaupart 2007). For the Archean, estimates of ambient T_p based on petrology vary according to the assumptions made and the method used ($\Delta T_p = +100$ – 250°C ; Herzberg et al. 2010, Condie et al. 2016, Putirka 2016, Ganne & Feng 2017, Aulbach & Arndt 2019), so although it is agreed that ambient mantle was warmer in the past, by how much remains an important unanswered question.

Hotter mantle has a lower viscosity, leading to more vigorous convection, and will melt to a greater extent, producing lithosphere with thicker primary (oceanic) crust—more magnesian than

WATER IN THE EARTH

An additional complication is the critical control of water on the viscosity of the mantle. Unfortunately, how water has been distributed between the interior and surface of Earth through time is uncertain. Some have argued that present-day Earth's mantle may be highly outgassed, containing only a small fraction of its original water content. Others have suggested that Earth in the Hadean had abundant surface water and a dry mantle, implying that regassing of the mantle has dominated Earth evolution. However, recent research has demonstrated that a deep hydrous mantle reservoir has been present in Earth's interior since at least the Paleoproterozoic. This is another unsolved and highly contentious issue that is fundamental to the question of whether plate tectonics could have operated on the Archean Earth, but one that is beyond our scope here.

present-day primary crust—above a more strongly depleted residue. Both the structure and composition of the lithosphere control its buoyancy, although not sufficiently to inhibit the viability of subduction during the Neoproterozoic (Korenaga 2013, 2017; Johnson et al. 2014; Weller et al. 2019). Subduction has been shown to be stable at ΔT_p up to $+100^\circ\text{C}$ if shear stress at the subduction interface is low (<5 MPa); if high (~ 30 MPa), subduction is stable only at present-day T_p (van Hunen & van den Berg 2008). At $\Delta T_p > +100^\circ\text{C}$, subduction is inferred to become increasingly less stable with earlier slab breakoff (van Hunen & van den Berg 2008, Sizova et al. 2010) (see the sidebar titled Water in the Earth).

We illustrate the ambiguity in resolving when plate tectonics could have emerged on Earth by reference to the thermal evolution of the mantle as modeled by Korenaga (2013, 2017) in relation to the T_p data from Herzberg et al. (2010). Assuming a low present-day Urey ratio (ratio of internal heat production to surface heat flux), the results show the T_p data are consistent with either a plate tectonic mode modeled with constant surface heat flow or a switch in heat-flow scaling from stagnant lid to plate tectonics at 3 Ga or 2 Ga, but not more recently.

2.3. Transitions Between Tectonic Modes

The critical parameters that control tectonic mode evolve with secular cooling—for example, internal energy decreases and lithospheric strength increases—potentially leading to a change in mode (Lenardic 2018). A transition from one mode to another—for instance, from stagnant to active lid—may take hundreds of millions of years (Bercovici & Ricard 2014). Alternatively, a planet could exhibit multiple tectonic modes in which the proportion of one relative to the other changes. For example, Earth currently has a bimodal distribution of plate sizes in which large plates are driven by the negative buoyancy of the subducting plate, whereas small plates are driven by convective tractions at their base (Hoink et al. 2013). With a higher mantle temperature and lower viscosity in the past, the sluggish lid component could have been larger such that the overall behavior was indistinguishable from a stagnant lid (Lenardic 2018). Turning the argument around, there could have been a smooth transition from a sluggish to an active lid mode with secular cooling (Foley 2018), consistent with the hypothesis that plate tectonics is part of an emergent, self-sustaining system as mantle convection evolves with time (Lenardic et al. 2019).

3. PLATE TECTONICS BACK TO THE ARCHEAN

3.1. How Do We Identify Plate Tectonics from the Geological Record?

There are clear similarities (e.g., Kusky et al. 2018) and differences (e.g., Gapais et al. 2009, Bédard 2018) between the geological features of Archean cratons and those of Paleoproterozoic

and younger orogenic belts. Critical to identifying a lower limit on the emergence of plate tectonics is the choice of geological information that may provide reliable evidence of a global network of narrow boundaries and multiple plates. Two types of information potentially fill this need: first, the demonstration of large, contemporaneous horizontal displacements between several continental fragments, consistent with multiple plates and the creation and destruction of oceanic lithosphere (e.g., Mitchell et al. 2014), and, second, the demonstration of widespread, contemporaneous bimodal metamorphism, which is the hallmark of asymmetric subduction at convergent plate boundaries at the present day (Brown 2006, Holder et al. 2019).

3.2. Supercontinents and Supercratons: Evidence for Large Horizontal Displacements

Based on widespread evidence of active convergent margins during the Proterozoic (e.g., Karlstrom et al. 2001, Giles et al. 2004, Roberts & Slagstad 2015, Pehrsson et al. 2016), we contend that plate tectonics has been continuous since its (last) emergence, as recorded by the supercontinent cycle (Nance et al. 2014). Proposals for plate tectonic shutdowns during the Proterozoic (Silver & Behn 2008, Stern 2018) may reflect a slower tempo (Sobolev & Brown 2019) or survivorship bias rather than a change in tectonic mode (see the sidebar titled The Supercontinent Cycle).

Seams in supercontinents are marked by orogenic belts, comprising variably deformed rocks with distinctive distributions of (mainly) marine facies sedimentary strata, igneous rocks, deformational structures, and metamorphic patterns that are commonly aligned subparallel to the belt (Cawood et al. 2009). These features are preserved during collisional orogenesis, marking where former ocean basins were eliminated. In the Phanerozoic and Proterozoic, abundant geological and paleomagnetic data provide evidence for large horizontal displacements since the beginning of the assembly of Columbia at c. 2.2 Ga (Mitchell et al. 2014, Li et al. 2016, Meert & Santosh 2017), in which at least 20 (of more than 30) major Archean cratons were ultimately sutured together by the Mesoproterozoic (Bleeker 2003, Pehrsson et al. 2013). However, before Columbia, large horizontal displacements are difficult to demonstrate globally. The cratons that formed Columbia were derived from the breakup of a few supercratons, which themselves formed during the late Mesoarchean to early Paleoproterozoic (Bleeker 2003, Pehrsson et al. 2013). The breakup is evidenced by late Neoarchean to early Paleoproterozoic mafic dike swarms associated with plumes (Bleeker & Ernst 2006, Ernst et al. 2013), and large horizontal displacements are supported by the abundance of passive margins at 2.2–1.8 Ga (Bradley 2008). Large horizontal displacements are also implied by the occurrence of a smaller number of passive margins at 2.7–2.4 Ga (Bradley 2008), but whether these reflect an active lid mode of global extent is uncertain.

THE SUPERCONTINENT CYCLE

The supercontinent cycle refers to the periodic aggregation of most of the extant continental lithosphere into a single entity. Several supercontinents are recognized: Pangea in the Mesozoic, Pannotia (including Gondwana) in the late Neoproterozoic, Rodinia in the early Neoproterozoic, and Columbia (including Nuna) in the Paleoproterozoic. In the Archean, fundamental differences between some of the better-known cratons suggest it is unlikely the extant continental lithosphere was ever aggregated into a single Archean supercontinent. Notwithstanding, most Archean cratons preserve evidence of rifted margins of Proterozoic age, which suggests they were once fragments of several larger entities or supercratons.

Were the number and size of the plates that existed in the Paleoproterozoic comparable to those that have existed since the breakup of Pangea? Based on results from three-dimensional spherical models of mantle convection, Mallard et al. (2016) proposed that, due to the feedback between mantle convection and lithosphere strength, the size distribution of plates has evolved with secular cooling of the mantle, such that Earth has fewer, larger plates now than earlier. This postulate supports a study suggesting that the global plate boundary network during the formation of Columbia was ~ 2.2 times the present length (Abbott & Menke 1990).

3.3. Indicators of Subduction in the Geological Record

Markers of subduction in the geological record are the most commonly used evidence of plate tectonics (Kusky et al. 2018). If widespread and contemporaneous, this evidence could be considered representative of a global network of plate boundaries (van Hunen & Moyen 2012). Widely used markers of subduction and collision are calc-alkaline igneous rocks, structural style in orogenic belts, and contemporaneous low-temperature–high-pressure metamorphism.

On contemporary Earth, igneous rocks generated above a subducting slab generally have particular trace element patterns (Pearce 2008), informally termed an arc signature. However, linking source, melting environment, and tectonic mode in variably metamorphosed igneous rocks preserving such a signature is highly contentious, particularly in the Archean (Pearce 2008, van Hunen & Moyen 2012, Bédard et al. 2013, Pearce 2014, Johnson et al. 2017, Smithies et al. 2018). Consequently, we do not rely on this signature as evidence of subduction, although its relevance in Archean rocks is discussed later.

Phanerozoic mountain belts include accretionary orogens related to continuous subduction in intraoceanic settings and at continental margins, and collisional orogens related to continent–continent or arc–continent collision that terminate subduction. The collisional type commonly overprints the accretionary type, creating complexity in the geological record. A common feature of collisional orogens is strong strain localization along lithospheric-scale thrusts that juxtapose units recording different types of metamorphism (Gapais et al. 2009). Similar features occur in some Neoproterozoic and Paleoproterozoic orogens (e.g., Weller & St-Onge 2017, Zibra et al. 2017), implying that the lithosphere was strong enough to allow modern-style tectonic processes. However, in Archean terrains tectonostratigraphic sequences in which older crustal units overlie younger crustal units across high-strain zones are rare. In southern West Greenland, examples include Mesoproterozoic–Neoproterozoic structures in the Nuuk region (Dziggel et al. 2014) and late Eoproterozoic structures in the Isua region (Nutman et al. 2013), although the latter example has been questioned (Webb et al. 2020). More commonly, the structural and metamorphic patterns of Archean and Proterozoic orogens imply distributed thickening and limited lateral variation in metamorphic grade (Gapais et al. 2009), consistent with a hotter lithosphere (Sizova et al. 2014).

The widespread occurrence of blueschists, low-temperature eclogites, and ultrahigh pressure metamorphic rocks in the geological record since the Neoproterozoic is unambiguous evidence of subduction (Brown 2006). However, these rock types are not ubiquitous in Phanerozoic orogenic belts, so their occurrence is not a requirement of subduction-driven collisional orogenesis, and they are scarce prior to the Neoproterozoic, suggesting their relatively recent appearance may be related to secular cooling (Brown 2006, van Hunen & van den Berg 2008, Sizova et al. 2014). Therefore, although metamorphism offers the possibility of tracing subduction back through time, we must first assess how metamorphism in collisional orogens changes with secular cooling of the mantle.

Information routinely retrieved from mineral assemblages in metamorphic rocks includes the temperature (T), pressure (P), and age of equilibration. Plotting the thermobaric ratio (T/P) against age permits evaluation of secular change in metamorphism (Brown & Johnson 2019a,c).

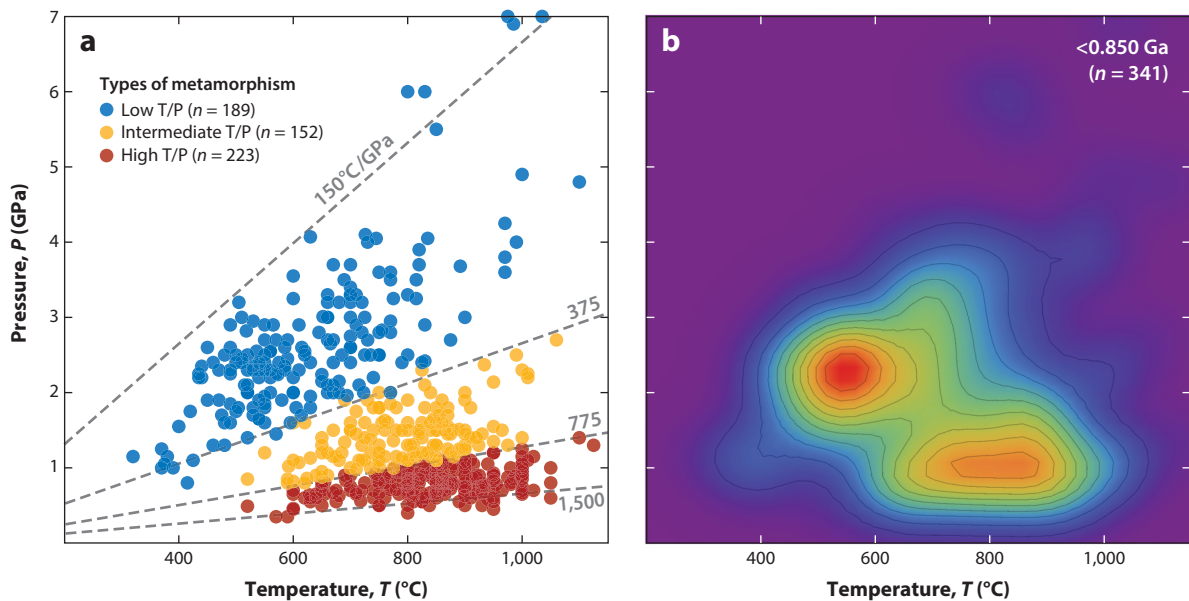


Figure 2

(a) Conditions of peak metamorphism for 564 localities with robust pressure (P), temperature (T), and age grouped by type, with representative thermal gradients (*thin gray dashed lines*). Three types of metamorphism are distinguished based on thermobaric ratios (T/P). (b) A plot of data <850 Ma in age contoured for density (contours are smooth kernel density estimates for which the bandwidth was computed automatically using DensityPlot in Wolfram Mathematica) to emphasize the characteristic bimodality of crustal metamorphism since 850 Ma. Panel *a* data from Brown & Johnson (2019b). Figure adapted with permission from Brown & Johnson (2019a).

Such data compiled for 564 localities from the Cenozoic to the Eoarchean (Brown & Johnson 2019b) define three types of metamorphism that are associated with particular tectonic settings on contemporary Earth (**Figure 2**): (a) low T/P metamorphism, comprising blueschists, and low-temperature and ultrahigh pressure eclogites, as found in subduction channels, accretionary orogens, and sutures in collisional orogens; (b) intermediate T/P metamorphism, comprising high-pressure granulites and medium-temperature and high-temperature eclogites, as found in collisional orogens; and (c) high T/P metamorphism, comprising migmatites and common and ultrahigh-temperature granulites, as found in orogenic hinterlands (orogenic plateaus or backarcs).

Based on the occurrence of intermediate and high T/P metamorphism, Brown & Johnson (2018; 2019a,c) argued that bimodal metamorphism can be recognized in the rock record since the beginning of the Neoarchean (**Figure 3a,b**). However, a statistical evaluation of T/P through time shows that global metamorphism >2.2 Ga can be fitted by a unimodal Gaussian distribution, whereas metamorphism ≤ 2.2 Ga is well described by bimodal mixed-Gaussian distributions (paired metamorphism) that become increasingly distinct with decreasing age (Holder et al. 2019). Thus, the Neoarchean might best be thought of as a period of transition between tectonic modes (cf. Cawood et al. 2018, Condie 2018), with the emergence of plate tectonics being complete by the early Paleoproterozoic.

During this transition, intermediate T/P metamorphism records subduction-driven collisional orogenesis, although maximum pressure was limited by shallow slab breakoff, whereas high T/P metamorphism records postbreakoff extension of the overriding plate in the orogenic hinterland

(Sizova et al. 2014, Chowdhury et al. 2017). The occurrence of subduction-driven collisional orogenesis in the Neoarchean is consistent with the sporadic occurrence of what, in seismic reflection profiles, are interpreted to be relict slabs beneath collisional belts in the Superior Craton (Percival et al. 2012). Alternatively, these features have been interpreted as due to overthrusting of oceanic crust by drifting proto-continents (Bédard 2018).

4. THE ARCHEAN ROCK RECORD

Plate tectonics has allowed us to make sense of almost all the large-scale features on Earth's surface. However, moving back in time, unambiguous recognition of volcanic arcs or their eroded remnants, or forearc sedimentary sequences, becomes increasingly difficult—particularly in the Archean, where field evidence is complex and ambiguous, and what remains of the crust is generally multiply deformed, fragmented, and metamorphosed. Because the rock record is limited and survivorship bias is a major concern, and Archean cratons generally do not preserve the uppermost levels of crust, where a record of subduction (for example, in the form of accretionary prisms and ophiolites) is more commonly preserved, evidence for the operation of plate tectonics in the Archean largely depends on geochemical indicators for subduction.

4.1. Survivorship Bias and the Geological Record in the Archean

Rocks older than Paleoarchean in age (>3.6 Ga) are rare. Furthermore, the age distribution of detrital zircon grains and magmatic rocks (e.g., Roberts & Spencer 2015, Johnson et al. 2019) shows a marked decline in the amount of preserved crust older than Neoarchean (>2.8 Ga). Thus, we may question to what extent preserved crust older than Neoarchean is representative, in terms of either relative volume of rock types or the processes it records. Is this crust the surviving remnants of a much larger volume that was destroyed (Hawkesworth et al. 2017, 2019; Dhuime et al. 2018, Korenaga 2018), or was there only ever a small volume produced (Condie et al. 2018)? How significant are the scarcity of Eoarchean crustal rocks and the almost complete absence of a Hadean rock record? In addition, there has been a strong geographic bias in studies of Archean crust, with the oldest known rocks attracting more attention. For these and other more practical reasons, much Archean and early Earth research has concentrated on a relatively small number of key localities.

4.2. Key Features of Archean Geology That Must Be Explained

Planetary differentiation is irreversible. Regardless of Earth's compositional and mineralogical structure immediately following the Moon-forming impact, the silicate Earth today is immeasurably more compositionally, mineralogically, and structurally complex. That Archean rocks are different from post-Archean geology is inevitable, irrespective of the tectonic mode in which they formed. With this in mind, we discuss some key features of Archean geology that warrant examination.

4.2.1. Lithological differences between pre- and post-Archean geology. Archean continental crust not only looks different, and was mapped as such by early geologists, but also is compositionally distinct from younger andesitic continental crust (Rudnick 1995). Furthermore, Archean upper crust may have been more mafic than post-Archean emergent crust (e.g., Taylor & McLennan 1985, Dhuime et al. 2015, Tang et al. 2016, Large et al. 2018, Chen et al. 2019).

Exposed Archean crust mostly comprises higher-grade gray gneiss terrains and lower-grade granite–greenstone belts with an archetypical dome-and-basin structure. These probably represent the lower and upper levels of the ancient continental crust, respectively (Johnson et al. 2016),

Figure 3 (Figure appears on preceding page)

(a) Metamorphic thermobaric ratios [temperature (T), pressure (P) (T/P)] for 564 localities grouped by type plotted against age. Three types of metamorphism are shown. The dashed lines show a second-order polynomial regression of the data for the high T/P (red) and a linear regression of the data for the intermediate T/P (orange) types, respectively. (b) T/P data for localities >850 Ma in age contoured for density (contours are smooth kernel density estimates for which the bandwidth was computed automatically using DensityPlot in Wolfram Mathematica) to emphasize the development of bimodality from the Neoproterozoic through the Proterozoic. (c) Histogram of ages and probability density function for metamorphism at 564 localities. Panels *a* and *c* adapted with permission from Brown & Johnson (2019a). Panel *b* adapted with permission from Brown & Johnson (2019c).

but they have also been interpreted as analogs of modern active continental arc margins and associated backarcs (Windley & Smith 1976).

Gray gneiss terrains are volumetrically dominated by sodic granitoids of the tonalite–trondhjemite–granodiorite (TTG) series. TTGs are compositionally similar to, but subtly different from, modern adakites that mostly form by partial melting of subducting hydrated oceanic crust (Smithies 2000, Condie 2005, Martin et al. 2005). Were TTGs also products of slab melting, or are other tectonic settings permissible? (See the sidebar titled The Tonalite–Trondhjemite–Granodiorite Suite of Rocks.)

In addition to felsic domes, granite–greenstone belts comprise sequences of mainly (metamorphosed) basalt, with subordinate ultramafic to felsic volcanic rocks and various sedimentary rocks. A significant feature is the common occurrence of komatiites, high-magnesian ($\text{MgO} > 18$ wt%) lavas that are rare in the post-Archean rock record. The unusually high mantle temperatures required to form komatiites (ΔT_p of $+200$ – 300°C ; Herzberg et al. 2007) are commonly explained by invoking more vigorous plume activity promoted by higher temperatures at the core–mantle boundary.

Most Archean lavas were erupted subaqueously, suggesting very little land had emerged before the late Archean–early Proterozoic (Flament et al. 2008, Rey & Coltice 2008, Bindeman et al. 2018). If the development of significant topography in the Proterozoic reflects strengthening of the continental lithosphere, was the Archean lithosphere too weak to focus deformation and break into plates?

4.2.2. Differences in the character of the Archean and post-Archean lithosphere. Crust that stabilized during the Mesoarchean–Neoproterozoic and has been largely undisturbed since is generally thinner than post-Archean crust and has fewer internal seismic discontinuities, slower seismic velocities in the lower part reflecting a more felsic composition, and a sharper, flatter Moho (Abbott et al. 2013). Given that Archean cratonic crust is dominated by TTGs mainly derived through partial melting of hydrated mafic sources, where is the (substantial) residual mafic lowermost crust and how and when was it lost? Could it have been returned to the mantle through

THE TONALITE–TRONDHJEMITE–GRANODIORITE SUITE OF ROCKS

Archean rocks are generally found in cratons, which form the stable interiors of continents. Archean cratons are dominantly composed of gneisses derived from a variety of initial rock types, but mostly of igneous origin, many of which have been tectonically transposed during strong deformation, obscuring their original relationships. Within these gneisses, an important group of igneous rocks is the tonalite–trondhjemite–granodiorite (TTG) suite, which is largely confined to the Archean. TTGs are silica-rich ($\text{SiO}_2 > 64$ wt%, but commonly ~ 70 wt% or greater) rocks with high Al_2O_3 (15.0–16.0 wt%) and Na_2O (3.0–7.0 wt%) contents, low $\text{K}_2\text{O}/\text{Na}_2\text{O}$ (< 0.5) ratios, and low total ferromagnesian oxide (Fe_2O_3 total + MgO + MnO + $\text{TiO}_2 \leq 5$ wt%) contents. TTGs generally have fractionated rare earth element (REE) patterns and low heavy REE contents.

density-driven delamination at a level corresponding to the garnet-in isograd (Abbott et al. 2013) or peeled off during subduction-driven collisional orogenesis on a hotter Earth (Chowdhury et al. 2017)?

Based on zircon ages and radiogenic isotopes, estimates of the overall volume of crust extracted from the mantle during the Hadean–Archean, the proportion of this crust returned to the mantle during the same period, and the volume of continental crust remaining at the surface at 3 Ga are equivocal (Dhuime et al. 2018, Korenaga 2018). In a recent study using molybdenum isotopes, McCoy-West et al. (2019) argued that, since 3.5 Ga, the composition of depleted mantle has essentially remained unchanged, implying that, since that time, the rates of crustal extraction and return have been in balance. If correct, such early depletion of the mantle requires extraction of a far greater volume of crust and higher rates of crustal growth and destruction prior to 3.5 Ga than generally expected. Dhuime et al. (2012, 2018) argued that 60–80% of the present volume of continental crust was present by 3 Ga. However, there is little if any evidence for the complementary mafic to ultramafic residues, which presumably were recycled back into the mantle (Abbott et al. 2013, Johnson et al. 2014, Chowdhury et al. 2017, Bédard 2018).

The thickness and composition of the continental lithospheric mantle (CLM) have changed through time (e.g., Pearson & Wittig 2014, Griffin & O'Reilly 2019). Archean CLM is thicker (up to ~250 km thick) and significantly more depleted (residual) than younger CLM, consistent with higher mantle temperatures and more extensive mantle melting (Lee 2003, Griffin et al. 2008). Importantly, the lower FeO content of more depleted Archean CLM means it has a lower density than younger CLM (e.g., Lee 2003, Johnson et al. 2014) and is buoyant with respect to the underlying asthenosphere, making it difficult to subduct or otherwise delaminate (Djomani et al. 2001, Lee et al. 2011). The Re–Os isotope composition of the oldest sulfide grains from mantle xenoliths, which likely date the timing of formation of the CLM, yields ages with a pronounced peak at c. 3.0–2.5 Ga (Pearson & Wittig 2008, Griffin & O'Reilly 2019), reflecting preservation of the oldest cratonic lithosphere during a transition to plate tectonics.

4.3. The Petrogenesis of Archean Crust

Modern arc basalts have characteristic compositions reflecting partial melting, on average at ~100 km depth below the active magmatic arc, of depleted mantle that was hydrated and enriched in incompatible elements through interaction with fluids derived from subducted materials (Pearce 2008). Relative to MORB, arc basalts are enriched in large ion lithophile elements (LILE); show fractionated rare earth element (REE) patterns, with preferential incorporation of light REE (LREE) over heavy REE (HREE); and are depleted in high field strength elements, leading to pronounced negative anomalies in Nb, Ta, Zr, Hf, and Ti. Although these characteristics are taken as a reliable proxy for subduction in many studies, they also occur in basalts derived from subduction-modified lithospheric mantle (Pearce 2008). However, the fundamental question here is whether modern arc basalts really do provide the key to the genesis of Archean basalts with similar compositional characteristics, or whether there are plausible mechanisms by which hydrated crust can be recycled into the mantle other than by subduction. Numerical modeling suggests that sagduction and dripping of the lowermost crust and crustal overturns may represent plausible alternatives (Johnson et al. 2014; Sizova et al. 2015; Fischer & Gerya 2016a, 2016b; Nebel et al. 2018; Roman & Arndt 2019).

The so-called arc signature is also found in most TTGs, which show extreme REE fractionation, with high chondrite-normalized La/Yb ratios ($La_N/Yb_N > 15$) and preferential enrichment of Sr over Y ($Sr_N/Y_N > \text{or} \gg 20$; Martin et al. 2014, Hoffmann et al. 2019). Notwithstanding that some TTGs may have formed through fractional crystallization (Liou & Guo 2019), the pronounced HREE and Y depletion in most TTGs (~80%; Moyen 2011) is interpreted to

record partial melting of hydrated basaltic rocks (amphibolite) at depth within the stability field of garnet. Although garnet may indicate pressures >1.5 GPa based on experimental studies of amphibolite (Moyen & Stevens 2006), its stability is highly sensitive to bulk rock Mg# [atomic $\text{Mg}/(\text{Mg} + \text{Fe}^{2+})$] and garnet may be stable at pressures as low as 0.7 GPa (Johnson et al. 2017). This lower pressure is within the range of thickness for primary crust generated during the Archean (Dhuime et al. 2015), and melting this crust may not require subduction (Sizova et al. 2015; also see Hoffmann et al. 2019, table 7.2).

Around 20% of TTGs record a high-pressure signature, considered to record extraction from an eclogite residue at pressure >2.0 GPa (Moyen 2011), corresponding to the extreme depths (>60 km) experienced by younger crustal rocks only during deep subduction (Brown & Johnson 2018). However, such high-pressure TTGs have recently been interpreted to represent fractionated sanukitoid (high-magnesian diorite) melts derived by partial melting of chemically enriched hydrated lithospheric mantle (Smithies et al. 2019).

5. COMPOSITIONAL CHARACTERISTICS OF THE ARCHEAN ROCK RECORD

5.1. The Eoarchean (4.0–3.6 Ga)

Earth's oldest known crustal rocks are 4.02–3.96 Ga magnetite-rich felsic gneisses that form a component of the Acasta Gneiss Complex, part of the Slave Craton in northwest Canada (Reimink et al. 2016, and references therein). These gneisses have compositions that are clearly distinct from TTGs and are considered to have formed through in situ partial melting of Fe-rich amphibolite at extremely low pressures, equating to the uppermost few kilometers of hydrated mafic crust (Johnson et al. 2018). Melting under such extreme thermal gradients is most easily explained through meteorite impact, which may have been the predominant mechanism for generating felsic rocks in the Hadean and early Eoarchean (Johnson et al. 2018). That the end of the purported late heavy bombardment (Koeberl 2006, Marchi et al. 2014) broadly coincides with survival of the oldest rocks in most cratons seems unlikely to be a coincidence. Furthermore, numerical simulations have shown that large bolide impacts may induce mantle upwellings capable of driving short-lived (>10 Ma) localized subduction events in a dominantly stagnant lid mode (O'Neill et al. 2017).

Earth's most extensive tract of Eoarchean crust ($\sim 3,000$ km²) is the 3.89–3.65 Ga Itsaq Gneiss Complex (IGC) in southwest Greenland (Hoffmann et al. 2014, Nutman & Bennett 2019), which is spatially associated with the 3.8–3.7 Ga Isua Supracrustal Belt (ISB) that includes amphibolites with tholeiitic and island arc affinities and even boninite-like compositions (Polat et al. 2002, Polat & Hofmann 2003, Jenner et al. 2009, Hoffmann et al. 2010). Together, the IGC and ISB form the Eoarchean core to the North Atlantic Craton—a collage of Eoarchean to Mesoarchean terranes defined on the basis of structural relationships, metamorphic grade, and protolith age—which was assembled during the Neoarchean (Friend & Nutman 2019, and references therein).

Southwest Greenland has become a classic model for terrane accretion during the Mesoarchean and Neoarchean, and subduction has been invoked to explain formation of even its earliest constituents (Polat et al. 2015, Windley & Garde 2009, Friend & Nutman 2019). On modern Earth, boninites are volcanic rocks associated with subduction initiation, and boninite-like rocks in the ISB have been argued to be a direct analog (Polat et al. 2002, Hoffmann et al. 2010). However, many Archean boninite-like rocks are silicious high-Mg, low-Ti basalts that could have formed in variably depleted plumes (Smithies et al. 2004, Pearce & Reagan 2019) or from otherwise anomalously hot upwelling mantle.

In **Figure 4** we show plots of Th/Yb versus Nb/Yb for basalts from various Archean localities, including those of Eoarchean and Mesoarchean age from southwest Greenland. This plot

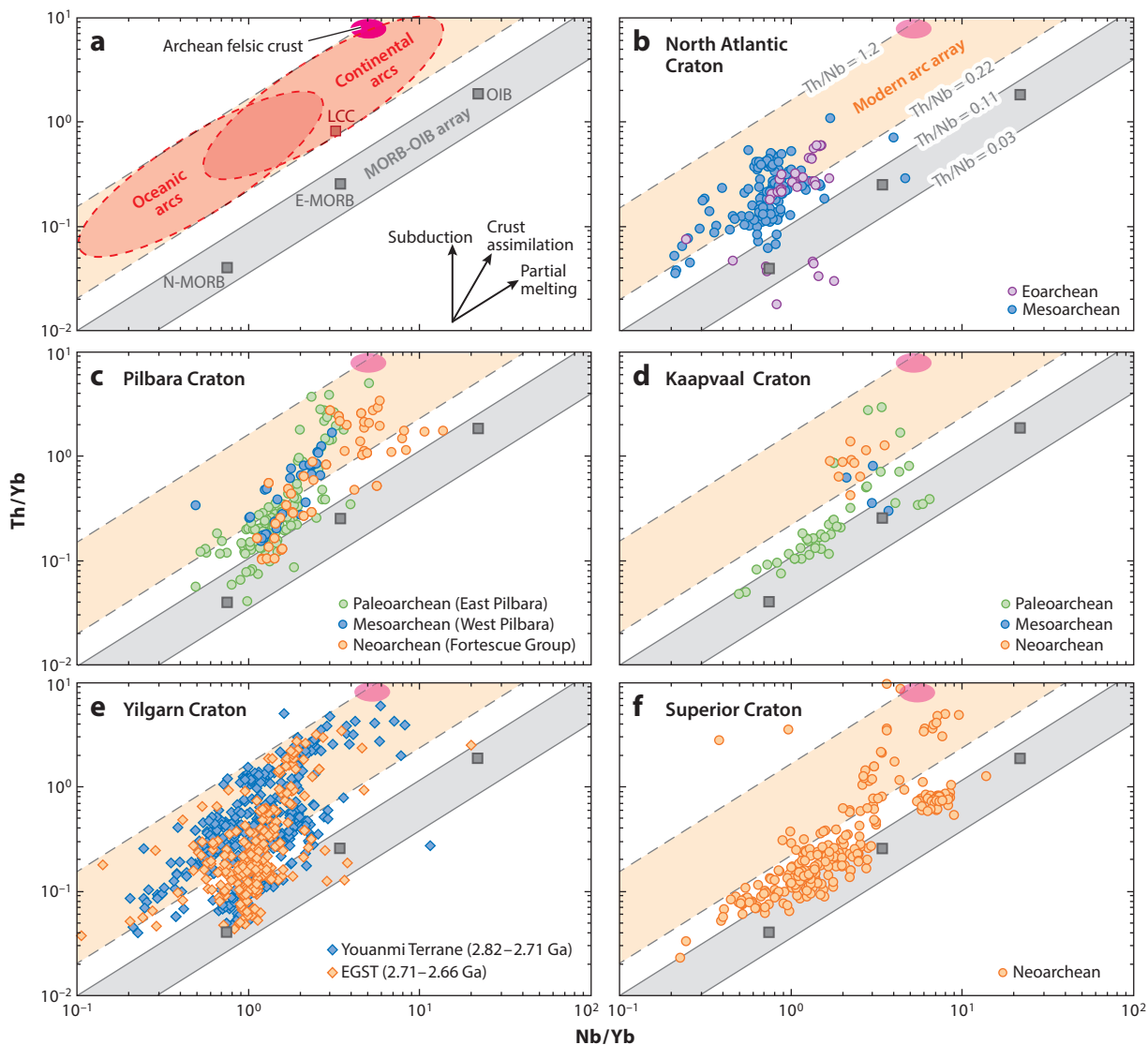


Figure 4

(a) The Th/Yb–Nb/Yb plot of Pearce (2008) showing main fields as well as composition vectors expected during subduction enrichment of a mantle source, crustal assimilation, and partial melting. Panel *a* adapted from Smithies et al. (2018). Copyright 2018, with permission from Elsevier; data for ocean floor basalt compositions (N-MORB, E-MORB, and OIB) from Sun & McDonough (1989); for the average composition of Archean felsic crust from Moyen (2011); and for modern LCC from Rudnick & Gao (2014). (b)–(f) Plots of Th/Yb versus Nb/Yb for basalts from the North Atlantic Craton (southwest Greenland), the Pilbara Craton, the Kaapvaal Craton (Barberton Granite–Greenstone Belt), and the Superior Craton (Abitibi Greenstone Belt) colored by age, and for the Yilgarn Craton colored by terrane. Data for panels *b*–*f* from GeoRoc (<http://georoc.mpch-mainz.gwdg.de/georoc/>) and GSWA (2019), filtered by MgO (<18 wt%) and SiO₂ (>45 and <57 wt%) content. Abbreviations: EGST, Eastern Goldfields Superterrane; E-MORB, enhanced mid-ocean ridge basalt; LCC, lower continental crust; MORB, mid-ocean ridge basalt; N-MORB, normal mid-ocean ridge basalt; OIB, ocean island basalt.

potentially allows distinction between basalts derived from mantle enriched in crustally derived Th—for example, by recycling through subduction—versus enrichment of basalts in both Th and Nb through crustal assimilation (Pearce 2008). Melting of a subduction-modified source and fractional crystallization generate an array of compositions with constant Th/Nb within the modern arc array (orange band, **Figure 4a**), which is parallel to but above the mid-ocean ridge basalt–ocean island basalt (MORB–OIB) array (gray band, **Figure 4a**). By contrast, mixing between MORB and a crustal component generates an array of compositions with variable Th/Nb that is steeper than either array and is unrelated to subduction (**Figure 4a**). Notwithstanding the difference in Th/Nb trends, caution should be exercised when attempting to fingerprint Archean tectonic settings (Pearce 2014).

Although the data are limited in number, the Th/Nb trends for Eoarchean basalts from southwest Greenland are offset from MORB and mostly plot within the modern arc array. Thus, these data appear to support mantle enrichment, which may be consistent with subduction-like processes (**Figure 4b**). Mantle enrichment in LREE through crustal recycling has also been proposed as the mechanism by which ϵHf was decoupled from ϵNd , as measured in some Eoarchean ISB tholeiites (Hoffmann et al. 2011). Modeling suggests an upper limit of 5% on the amount of crustal contaminant, consistent with the extracted basalts and TTG crust (Gardiner et al. 2019a). However, others propose that such isotopic decoupling is an artifact of later LREE mobilization during metamorphism (Kemp et al. 2019) and is not indicative of subduction. In summary, although there is strong evidence for a recycled component in the source of Eoarchean mafic rocks from southwest Greenland, the tectonic process by which this occurred remains ambiguous.

5.2. The Paleoproterozoic (3.6–3.2 Ga)

The East Pilbara Terrane (EPT) in Western Australia and the Barberton Granite–Greenstone Belt (BGGB) in the Kaapvaal Craton of southern Africa are among the most intensively studied Archean granite–greenstone terranes.

The EPT forms the Paleoproterozoic core to the Pilbara Craton and has a well-developed dome-and-keel architecture comprising steeply dipping supracrustal sequences surrounding domal granitoid complexes. A mafic substrate is proposed to have formed through episodic volcanic resurfacing from 3.8 to 3.5 Ga until it reached a critical depth (>30 km) allowing intracrustal melting to produce TTG magmas (Van Kranendonk et al. 2007). Composite TTGs form domes developed via punctuated magmatic events from c. 3.5 to 3.2 Ga (Smithies et al. 2005b). Hf and Nd isotopes measured in the 3.5–3.2 Ga TTG suites show an increase in unradiogenic ϵHf and ϵNd with time, with consistent model ages of c. 3.8–3.7 Ga (Smithies et al. 2003, Gardiner et al. 2017), implying derivation from an Eoarchean mafic source. These trends support a tectonic model in which sagduction episodically delivered (hydrated) mafic volcanic rocks into a deep melting zone largely comprising an older residual mafic substrate.

A petrogenetic model for the EPT must account for the TTGs being too enriched in LILE and LREE to have been derived from a primary mantle melt. Their source was likely an enriched basalt—either the product of multistage crustal reworking (Johnson et al. 2017) or extracted at c. 3.5 Ga from a mantle pre-enriched through recycling of a now-absent c. 3.7 Ga felsic crust (Smithies et al. 2009). The trace element composition of the TTGs requires 20–30% melting of this enriched source at thermal gradients much warmer than those associated with modern subduction (Johnson et al. 2017), suggesting an alternative tectonic setting. The existence of earlier felsic crust is supported by the detrital zircon Hf isotope record, which extends back to at least 3.65 Ga (Gardiner et al. 2019c, and references therein), and the variable Th/Nb trends of Pilbara basalts (**Figure 4c**), which are consistent with crustal assimilation rather than derivation from mantle enriched through subduction.

Paleoarchean domal granitoid complexes in the BGGB have been compared to those in the EPT in terms of a volcanic plateau origin (e.g., Van Kranendonk et al. 2015), but they differ in detail. Generally the BGGB granites are not polychronic composite intrusions, and where they comprise more than one component, these have similar magmatic ages (e.g., Kamo & Davis 1994). Based on Hf isotopes, TTGs that crystallized at c. 3.2 Ga in the northern Barberton region were derived from a predominantly juvenile source, distinguishing them from an older suite of TTGs (c. 3.5–3.4 Ga) in the southern Barberton region that exhibits less radiogenic Hf isotope values, implying some crustal assimilation (Zeh et al. 2009). However, in general TTGs in the BGGB are less enriched than those from the EPT (e.g., Moyén et al. 2019), suggesting crustal assimilation was limited. Limited crustal assimilation is also implied by the Th/Yb versus Nb/Yb plot, where Paleoarchean basalts largely follow constant Th/Nb trends with slightly elevated values at higher Nb/Yb (**Figure 4d**). It is unclear whether these distinctions between the EPT and the BGGB genuinely reflect formation within different tectonic settings (i.e., intraplate versus subduction; Kisters et al. 2010, and references therein; Schneider et al. 2019, and references therein) or simply a different prehistory of crustal recycling.

5.3. The Mesoarchean (3.2–2.8 Ga) and Neoarchean (2.8–2.5 Ga)

The late Archean was a crucial period for potential changes in tectonic mode because it witnessed shifts in various proxies that have been linked to the emergence of plate tectonics (e.g., Taylor & McLennan 1985; Condie 1993; Dhuime et al. 2012, 2015; Næraa et al. 2012; Tang et al. 2016; Greber et al. 2017; Large et al. 2018; Moyén & Laurent 2018; Nebel-Jacobsen et al. 2018; Chen et al. 2019; Greber & Dauphas 2019; Johnson et al. 2019; Ravindran et al. 2020). However, the interpretation of proxies reminds us of the parable of the blind men and an elephant. Studies based on proxies have obvious merit, but do they all address the problem as a whole? Are some of the so-called global data sets really only local—for example, due to survivorship bias or the late Archean rise of the continents—and do some data sets record an increased proportion of subduction (e.g., in a sluggish lid mode) rather than a change to plate tectonics? Does the spread of ages centered around c. 3.0 Ga that has been argued to record the emergence of plate tectonics simply reflect the beginning of a long transition from a sluggish lid to plate tectonics? Certainly, a change around 3.0 Ga is not clearly recorded by basalts from several cratons, including Pilbara, Yilgarn, and Superior (Smithies et al. 2018; cf. Moyén & van Hunen 2012), where the record appears to vary between magmatism with arc and nonarc characteristics (**Figure 4**). Maybe those proxies that are global record the onset of the transition, and differences between cratons reflect the nature of the transition.

In the Pilbara Craton, a change in tectonic style from sagduction to subduction and terrane accretion has been proposed at c. 3.2–3.0 Ga (e.g., Hickman 2004, Smithies et al. 2007) and linked with an inferred Wilson cycle and craton assembly at c. 3.0 Ga (Van Kranendonk et al. 2010). Evidence for this transition includes a shift toward more juvenile Nd isotope compositions after c. 3.2 Ga and lithological evidence of subduction-related magmatism, including sanukitoids, Nb-enriched tholeiites, and adakites (Smithies et al. 2005a). Nevertheless, an investigation of secular change in the composition of Pilbara basalts using variation in Th/Yb versus Nb/Yb identified two contrasting trends (Smithies et al. 2018). Basalts with constant Th/Nb trends that might imply melting of a subduction modified source are only seen in the Mesoarchean (c. 3.13–2.95 Ga in the West Pilbara Terrane; **Figure 4c**), whereas most of the older (Paleoarchean of the EPT) and younger (Neoarchean of the Fortescue Group) basalt sequences have variable Th/Nb trends that do not appear subduction-like (**Figure 4c**).

Much of the Yilgarn Craton formed between c. 3.1 and 2.6 Ga. Magmatic rocks vary in composition from TTG to monzogranite with decreasing age and commonly record geological

and geochemical characteristics considered consistent with subduction, including the presence of sanukitoids. In the Th/Yb versus Nb/Yb plot (**Figure 4e**), c. 2.82–2.71 Ga basalts from the western Youanmi Terrane define subduction-like trends, whereas c. 2.71–2.66 Ga basalts from the Eastern Goldfields Superterrane do not (Smithies et al. 2018). These data support the idea of short-lived episodes of subduction that may not record plate tectonics (O'Neill et al. 2018) but may be characteristic of the transition to plate tectonics (O'Neill et al. 2019).

We note that subduction-related rocks such as sanukitoids occur in many late Archean terranes, but are not present in pre-3.0 Ga crust (Smithies 2000). Furthermore, statistical interrogation of large geochemical and isotopic data sets has identified shifts during this period that may reflect changing tectonic modes. According to Condie (2018), the proportion of basalts derived from hydrated mantle sources (i.e., arc-like) increased significantly during the Mesoarchean–Neoarchean, interpreted as a transition to plate tectonics. There was also an increasing proportion of basalts derived from clearly depleted or enriched mantle sources from the Archean to the Proterozoic, suggesting that the distinction between magma sources was less significant during the Archean (Condie 2018, Moyen & Laurent 2018). However, using weighted bootstrap resampling to minimize any preservation and/or sampling bias, Keller & Schoene (2018) found that basalts show a gradational change through the Archean, consistent with secular cooling, but with no requirement for a sharp change in tectonic mode. This type of change may be consistent with an evolution from a sluggish lid to an active lid tectonic mode.

During the Mesoarchean–Neoarchean, a secular evolution of magma type from early TTGs to sanukitoids to potassic two-mica granites is recognized in many cratons, albeit at different times (Cawood et al. 2018). Furthermore, the chemical composition of TTGs shows an increasing spread through the Archean, reflecting the irreversible differentiation of the lithosphere (Gardiner et al. 2017, Johnson et al. 2019). Despite a high degree of scatter, the ratios of several important geochemical proxies—particularly K_2O/Na_2O , Sr/Y, and La_N/Yb_N —show statistically significant change in the interval 3.3–3.0 Ga (Johnson et al. 2019) that may identify the appearance of continental arcs and the beginning of the transition to plate tectonics.

In summary, there is geochemical evidence for subduction initiation and periods of stable subduction at multiple localities between c. 3.2 and 2.5 Ga. This period also marks the assembly of proto-continents into supercratons via accretion of proto-continents (e.g., Bleeker 2003, Smithies et al. 2007, Percival et al. 2012, Cawood et al. 2018, Friend & Nutman 2019). We interpret these data as recording a transitional period that ultimately led to plate tectonics.

6. DISCUSSION: SUBDUCTION AND PLATE TECTONICS IN THE ARCHEAN

If we accept a Columbia supercontinent composed of Archean cratons with passive margin sequences that were sutured by orogenic belts, the earliest of several supercontinent configurations, then Earth has operated according to plate tectonic theory since at least the early Paleoproterozoic. To assess what tectonic mode preceded plate tectonics, we refer to numerical modeling studies of tectonics in the Archean so that we may evaluate them against the geological record (Gerya 2014, 2019) (see the sidebar titled Numerical Modeling).

6.1. What Preceded Plate Tectonics?

A prerequisite for plate tectonics is the ability to break the lithospheric lid and initiate plate boundaries. With increasing temperature, the viscosity of the mantle is lowered, such that viscous coupling between the convecting mantle and the overlying lithosphere is reduced. In mantle

NUMERICAL MODELING

Over the past two decades, much of the progress in evaluating plausible tectonic modes for early Earth has stemmed from numerical modeling. The exponential rise in computer processing power has permitted consideration of increasingly complex rheological and material properties that more closely simulate nature, using input criteria based on existing geological, geochemical, geophysical, and petrological data. This has allowed exploration and quantitative testing of various geodynamic hypotheses. Of particular importance to the input parameters is the development of thermodynamic models appropriate to partial melting of a wide spectrum of compositions ranging from ultrabasic to granitic. These advances allow calculation of the changes in phase assemblage, the abundance and composition of phases, and the densities of phase assemblages in crust–mantle systems that include solid, melt, and volatile phases, all as a function of pressure, temperature, and protolith composition.

convection models that use a pseudoplastic rheology for the lithosphere, higher T_p favors a stagnant lid for the early Earth (e.g., O'Neill et al. 2016). By contrast, models in which plate boundaries develop by grain size reduction favor a sluggish lid for the early Earth (e.g., Foley 2018). Higher temperature also weakens the slab, leading to more frequent slab breakoffs (van Hunen & van den Berg 2008) and, at ΔT_p around $+175^\circ\text{C}$, subduction may no longer be possible (Sizova et al. 2010). At ΔT_p higher than $+200^\circ\text{C}$, foundering by Rayleigh–Taylor instabilities (drips) of overthickened crust (Johnson et al. 2014, Nebel et al. 2018) is likely. Thus, a stagnant or sluggish lid mode, with a deformable (squishy) lithosphere (Rozel et al. 2017) punctuated by episodes of short-lived subduction (Sizova et al. 2015), may have preceded plate tectonics. The earlier mode may have been linked to plate tectonics via an episodic lid mode (O'Neill et al. 2016, 2018), which may be reflected in changes from arc to nonarc magmatism in the Mesoarchean–Neoproterozoic rock record of some cratons.

6.2. Formation and Breakup of Supercratons: The Transition to Plate Tectonics in the Neoproterozoic–Early Paleoproterozoic

Age-constrained geological and geochemical data retrieved from cratons demonstrate differences in petrogenesis and tectonic style from one to another without a uniform pattern. Prior to the Mesoarchean, the formation and evolution of some Archean crust (e.g., the North Atlantic Craton) may have been related to subduction-driven tectonics, whereas the development of other Archean crust (e.g., the EPT in the Pilbara Craton) appears to have been related to intraplate activity. Such differences may also occur between terranes within the same craton. For example the Mesoarchean Akia Terrane, now part of the North Atlantic Craton, may have formed in a stagnant lid mode (Gardiner et al. 2019b), whereas the Eoarchean IGC may have had a subduction origin. Potentially, these differences can be accommodated in more than one tectonic mode, such as sluggish lid (Foley 2018) or lid and plate tectonics (Capitanio et al. 2019).

In a sluggish lid mode, assuming weak plate boundaries can form during mantle convection at Archean conditions (Foley 2018), subduction may occur on a locally confined scale associated with lithospheric mobility, whereas intraplate magmatism may occur elsewhere due to passive mantle upwelling. On secular cooling, as subduction became more widespread, intraplate magmatism waned, and subduction-related magmatism increased during a transition to an active lid mode (Foley 2018). However, it remains unclear when this transition began or was completed. The beginning of such a transition may be what is recorded in the geology of the North Atlantic Craton and in the BGGC during the Mesoarchean–Neoproterozoic. By contrast, during the same period in other cratons (Pilbara, Yilgarn, and Superior), the evidence suggests only short-lived

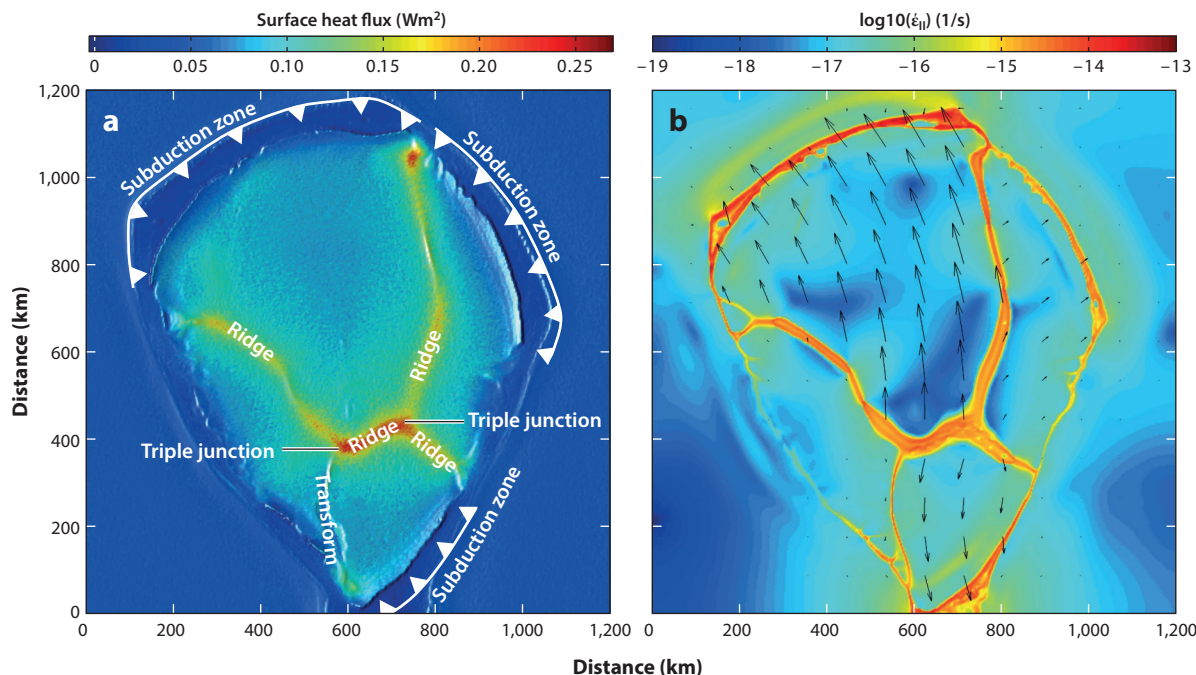


Figure 5

Development of an embryonic mosaic of plates separated by spreading centers (ridges), triple junctions, and transform faults at the latest stage of plume-induced subduction. (a) Surface heat fluxes projected onto the modeled surface topography, showing a pattern of spreading centers (*white lines with triangles* show dip directions of retreating subducting slabs). (b) Spatial distribution of the second strain rate invariant at a depth of 20 km (*arrows* show horizontal velocities of individual, young, nonsubducting plates moving toward retreating subducting slabs). Figure adapted from Gerya et al. (2015); copyright 2015, with permission from Springer Nature.

episodes of subduction that cannot necessarily be used to argue for plate tectonics (O'Neill et al. 2018, 2019). Thus, the Mesoarchean–Neoproterozoic may register the mixed signals we might expect of a transition to plate tectonics that was not completed until the early Paleoproterozoic.

How the transition to plate tectonics occurred remains cryptic, but during converging mantle flow a general requirement for initiation of subduction is thought to be the development of heterogeneities in rheology, buoyancy, and composition along the margins of early-formed protocontinents (Bercovici & Ricard 2014, Rey et al. 2014). In addition, other mechanisms may have contributed to the emergence of plate tectonics within a sluggish lid tectonic mode, such as subduction induced through intraplate plume activity (Gerya et al. 2015) (**Figure 5**). In the Archean, plume activity could have led to the formation of retreating subduction zones similar to those associated with the Caribbean large igneous province in the Cretaceous (Gerya et al. 2015). At higher ΔT_P of $+200^\circ\text{C}$, models of plume–lithosphere interactions show that plume-induced subduction is feasible for older lithosphere (80 Myr), whereas plume interaction with younger lithosphere (20 Myr) favors episodic dripping rather than stable subduction (Gerya et al. 2015) (**Figure 6**).

According to Perchuk et al. (2019), at higher ΔT_P of $+150^\circ\text{C}$, retreating subduction, as occurs in the plume model, could lead to an initial arc-free regime as formation of magma by fluid-fluxed melting in the wedge is depressed during rapid trench rollback and vigorous upwelling of asthenospheric mantle. Instead, decompression melting of the upwelling asthenospheric mantle results in widespread development of voluminous hydrous intraplate basalts (Perchuk et al. 2019). In this context, TTG crust could be generated by partial melting of the intraplate basalts

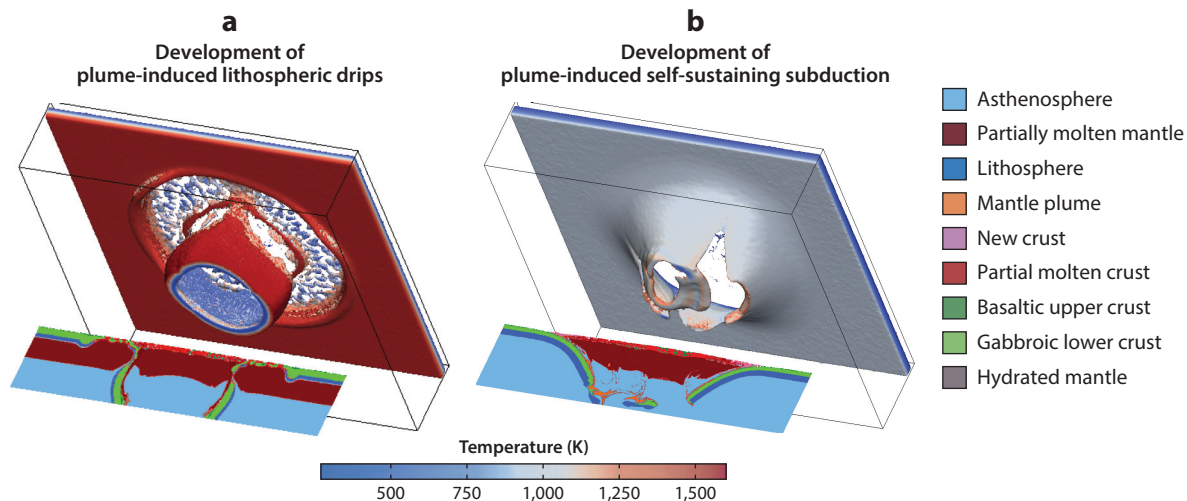


Figure 6

Plume–lithosphere interaction for hotter mantle temperature (ΔT_p of $+200^\circ\text{C}$) and thicker oceanic crust. (a) Development of plume-induced lithospheric drips for 20-Myr oceanic plate with 30-km-thick crust. (b) Development of plume-induced self-sustaining subduction for 80-Myr oceanic plate with 20-km-thick crust. Figure adapted from Gerya et al. (2015); copyright 2015, with permission from Springer Nature.

(Johnson et al. 2014, 2017; Sobolev & Brown 2019) (**Figure 7a**). Collision between these nascent proto-continental crustal domains and thickening of the lithosphere to create proto-continentals could have occurred as mantle downwelling eliminated the intervening ocean between a pair of retreating subduction cells (cf. Wang et al. 2018, Tang et al. 2019) (**Figure 7b**).

According to Sobolev & Brown (2019), the emergence and evolution of plate tectonics on Earth could have been related to the rise of the continents (Flament et al. 2008, Rey & Coltice 2008) and the accumulation of sediments at continental edges, and subsequently in trenches, to lubricate and stabilize subduction. In southern Africa, glaciation is recorded in the Mesoarchean Mozaan Group (c. 2.9 Ga; the upper part of the Pongola Supergroup) on the passive margin of the Kaapvaal Craton (Young et al. 1998). This combination of events may have allowed a spreading proto-continent to initiate subduction at its edge, as proposed by Rey et al. (2014) and documented for young subduction zones near Sulawesi in the Southwest Pacific (Hall 2019). Therefore, an early record of plate tectonics in the amalgamation of the Kaapvaal Craton is not surprising (Shirey & Richardson 2011).

We posit that the emergence of plate tectonics on Earth could be related, at least in part, to the development of regional networks of plate boundaries formed by intraplate plume-induced subduction initiation, formation of proto-continentals, and initiation of subduction at the edges of these proto-continentals. Thus, as a sluggish lid evolved toward an active lid, plume-induced regional networks could have linked with existing subduction sites gradually developing a global network of plate boundaries. During this late Mesoarchean to early Paleoproterozoic transition to plate tectonics, as secular cooling of the mantle became dominant over heating, proto-continental lithosphere was aggregated into a small number of supercratons. This process took hundreds of millions of years to accomplish, and there is no evidence of a single global supercraton-forming event. Subsequently, the supercratons were fragmented in response to mantle insulation (Lenardic 2017, and references therein) and/or possibly by plumes triggered by subduction around the supercraton margins.

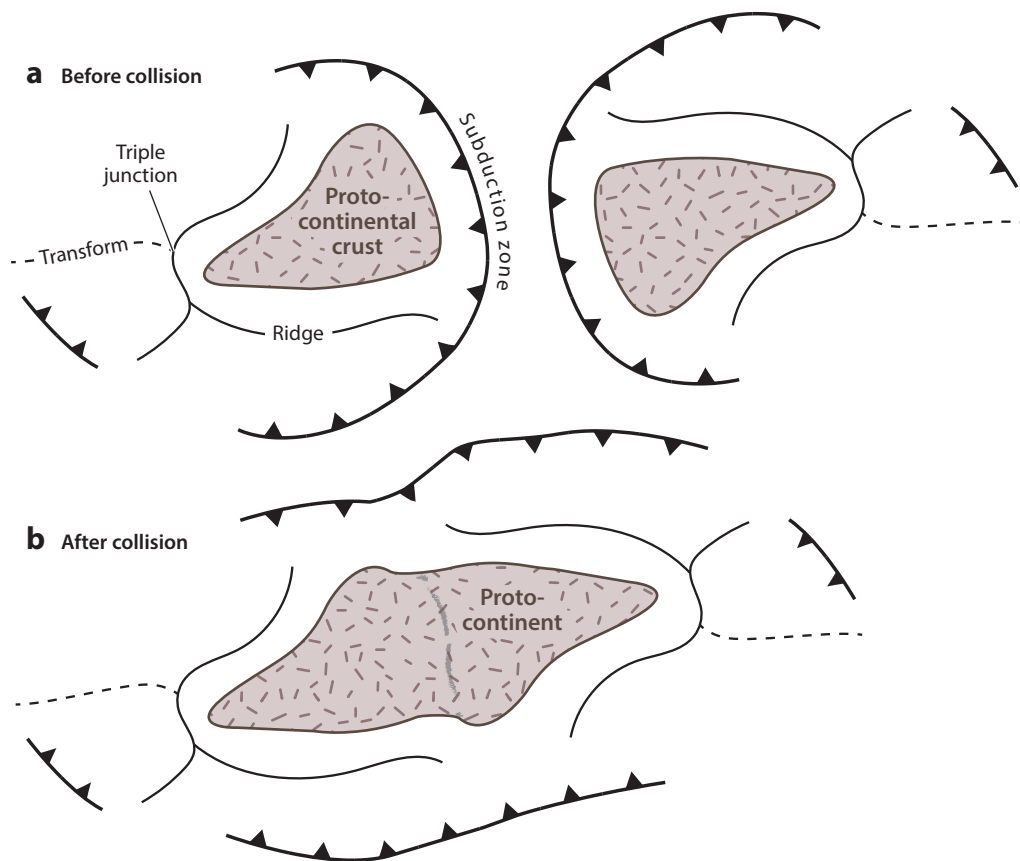


Figure 7

Sketches showing approaching plume-induced subduction cells (*a*) before elimination of the intervening oceanic lithosphere and collision and (*b*) after collision and formation of a proto-continent.

SUMMARY POINTS

1. Before the Mesoarchean–Neoproterozoic, a stagnant lid or sluggish lid or lid and plate tectonics mode, with a deformable (squishy) lithosphere and either intermittent (unstable) subduction or subduction on a locally confined scale, may have been dominant. A hotter mantle prohibited widespread stable subduction and plate tectonics, and magmatism was largely due to upwelling mantle and plume–lithosphere interactions.
2. During a Mesoarchean–Neoproterozoic transition, as secular cooling overwhelmed heat production, plate tectonics emerged, possibly in part via plume-initiated retreating subduction cells that generated proto-continents.
3. The rise of proto-continents and enhanced erosion to provide sediments at their edges may have enabled initiation of subduction under the proto-continents by spreading of continental margins over oceanic lithosphere. This process could have stabilized subduction and promoted the spread of plate tectonics globally, as recorded by the

widespread appearance of high and intermediate T/P metamorphism in the Neoarchean. The formation and breakup of the supercratons, and the formation of the first supercontinent Columbia, demonstrate that plate tectonics was fully developed by the early Paleoproterozoic.

FUTURE ISSUES

1. Geochemical criteria need to be identified to distinguish impact-induced from plume-induced melt products, so that we may determine the contribution of each process to the formation of the Hadean and Eoarchean tonalite-trondhjemite-granodiorite crust.
2. To evaluate hypotheses regarding the tectonics of early Earth, it is necessary to increase the number of metamorphic pressure, temperature, and age data from exposures of ancient (older than 2.8 Ga) continental crust worldwide. Retrieval of such data should be a priority of metamorphic studies.
3. To understand secular change in tectonics on Earth requires a better knowledge of the thermal evolution of Earth and mantle potential temperature and further research to resolve the contentious issue of the amount and temporal evolution of water in the mantle.
4. To understand the evolution of life on Earth, it is necessary to assess linkages between secular change in tectonic mode and the evolution of Earth's atmosphere, oceans, and landscape, since these control nutrient supply and environment.

DISCLOSURE STATEMENT

The authors are not aware of any affiliations, memberships, funding, or financial holdings that might be perceived as affecting the objectivity of this review.

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By necessity, in attempting to cover such a broad and (currently) popular topic, we have been able to cite only a small proportion of the relevant published literature. In most instances we have cited recent contributions, such that the interested reader may more easily follow the historical trail of ideas pertaining to Archean tectonics. T.J. acknowledges support from the State Key Laboratory for Geological Processes and Mineral Resources, China University of Geosciences, Wuhan (Open Funds GPMR201704 and GPMR201903). N.J.G. acknowledges Australian Research Council grant FL160100168 for financial support. We thank Dr. Marzieh Baes of the Helmholtz Centre Potsdam for provision of vector files that we modified in producing **Figures 5** and **6**. We are grateful to P. Cawood, C. Hawkesworth, and R. Rudnick for their comments that undoubtedly improved the published version of the paper. Any remaining errors or misconceptions are ours.

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